

Simulating the EOR with self-consistent growth of galaxies

Master's thesis presentation

ETH Zürich, University of Zürich

Simulating the Epoch of Reionization

BEoRN

Halo growth

Implementation

Results

End

Bibliography

Simulating the Epoch of Reionization

- The Epoch of Reionization
- The 21-cm signal
- Expression the 21-cm signal
- The current state of simulations
- Matrix
- Contained
- Four columns
- Two columns

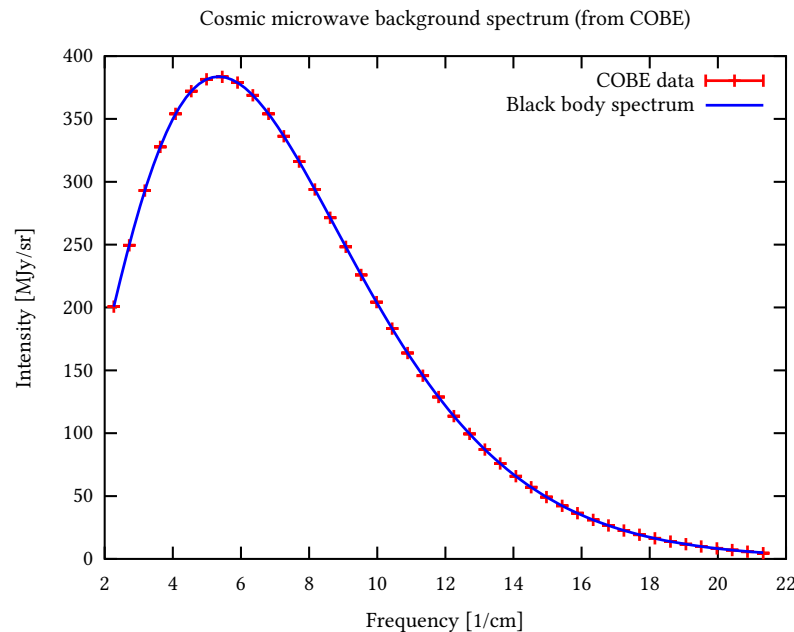
The 1.1 GHz 21-cm signal

- Marks the universe's last major phase transition: from neutral to ionized hydrogen.
- Shapes the large-scale structure of the intergalactic medium (IGM).
- Is strongly linked to the formation and growth of the first galaxies and black holes.
- Sets the stage for many observables: CMB secondary anisotropies, 21-cm signal, high- z galaxy surveys.

The 21-cm signal

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The brightness temperature describes the difference between the CMB temperature and the spin temperature of neutral hydrogen

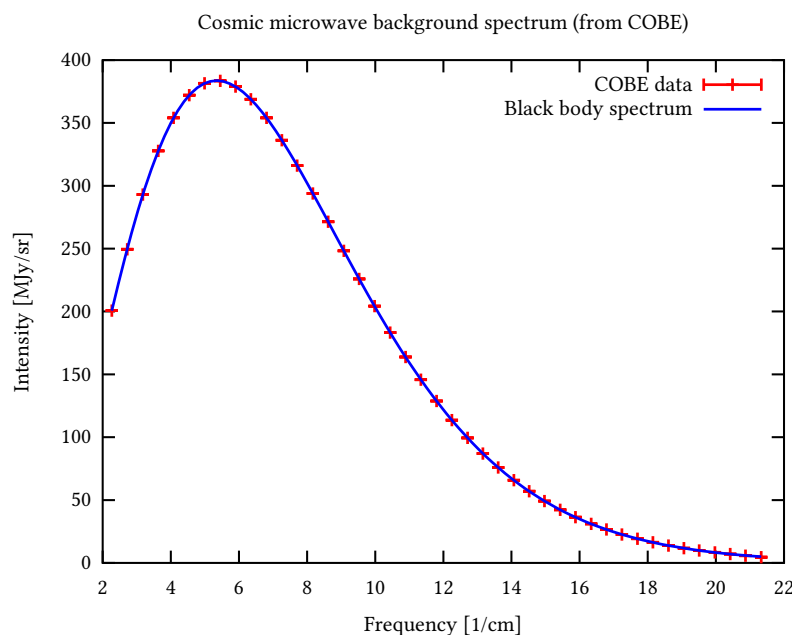


from [1]

The 21-cm signal

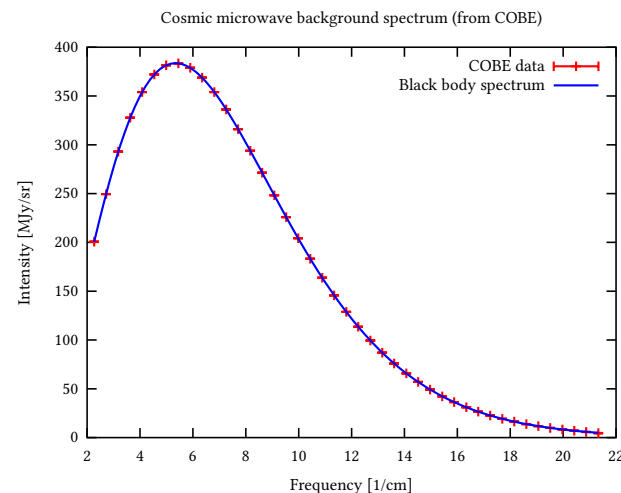
3 / 35

The brightness temperature describes the difference between the CMB temperature and the spin temperature of neutral hydrogen



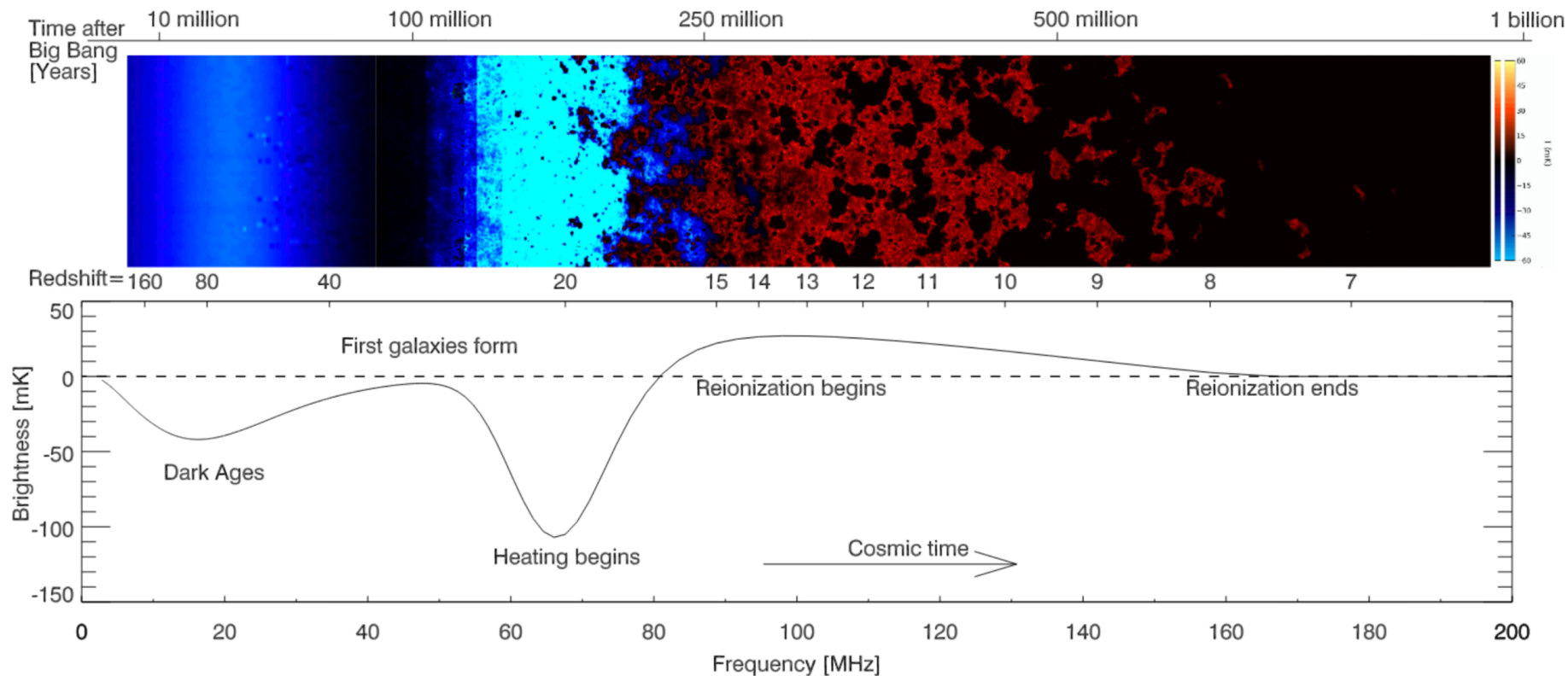
from [1]

Removing the contribution from the black body spectrum of the CMB yields the explicit 21-cm signal:



The 21-cm signal

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from [2]

Expressing the brightness temperature (e.g [2]):

$$dT_b(\boldsymbol{x}, z) \simeq T_0(z) \cdot x_{\text{HI}}(\boldsymbol{x}, z) \cdot (1 + \delta_b(\boldsymbol{x}, z)) \cdot \frac{x_\alpha(\boldsymbol{x}, z)}{1 + x_\alpha(\boldsymbol{x}, z)} \cdot \left(\frac{1 - T_{\text{CMB}}(z)}{T_{\text{gas}}(\boldsymbol{x}, z)} \right)$$

- further modulation by *RSD*

Traditional approaches:

→ slow and big

→ radiative transfer

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→ slow and big

→ radiative transfer

⇒ semi-numerical approaches such as
BEoRN [3]

- fast
- flexible

⇒ repeatable

Lineal

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Lineal

Following [4], [5]:

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$$\rho_{\alpha}(r \mid M, z) = \frac{(1+z)^2}{4\pi r^2} \cdot \sum_{n=2}^{n_m} f_n \cdot \varepsilon_{\alpha}(\nu') \cdot f_{\star} \cdot \dot{M}(z' \mid M, z)$$

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$$\frac{3}{2} \cdot \frac{d\rho_h(r \mid M, z)}{dz} = \frac{3\rho_h(r \mid M, z)}{1+z} - \frac{\rho_{\text{xray}}(r \mid M, z)}{k_B(1+z)H(z)}$$

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$$\frac{dV_b}{dt} = \frac{\dot{N}_{\text{ion}}(t)}{\bar{n}_H^0} - \alpha_B \cdot \frac{C}{a^3} \cdot \bar{n}_H^0 \cdot V_b$$

$$dT_b(\boldsymbol{x}, z) \simeq T_0(z) \cdot x_{\text{HI}}(\boldsymbol{x}, z) \cdot (1 + \delta_b(\boldsymbol{x}, z)) \cdot \frac{x_\alpha(\boldsymbol{x}, z)}{1 + x_\alpha(\boldsymbol{x}, z)} \cdot \left(\frac{1 - T_{\text{CMB}}(z)}{T_{\text{gas}}(\boldsymbol{x}, z)} \right)$$

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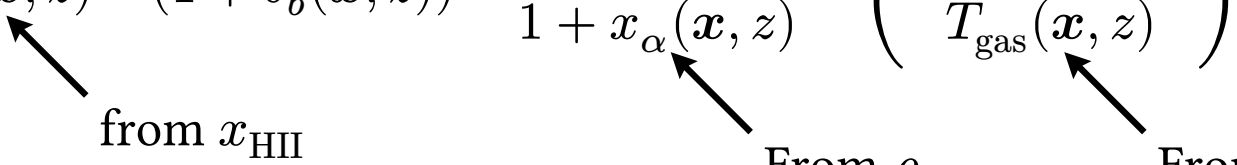
from x_{HII}

$$dT_b(\boldsymbol{x}, z) \simeq T_0(z) \cdot x_{\text{HI}}(\boldsymbol{x}, z) \cdot (1 + \delta_b(\boldsymbol{x}, z)) \cdot \frac{x_\alpha(\boldsymbol{x}, z)}{1 + x_\alpha(\boldsymbol{x}, z)} \cdot \left(\frac{1 - T_{\text{CMB}}(z)}{T_{\text{gas}}(\boldsymbol{x}, z)} \right)$$



from x_{HII} From ρ_α

$$dT_b(\mathbf{x}, z) \simeq T_0(z) \cdot x_{\text{HI}}(\mathbf{x}, z) \cdot (1 + \delta_b(\mathbf{x}, z)) \cdot \frac{x_\alpha(\mathbf{x}, z)}{1 + x_\alpha(\mathbf{x}, z)} \cdot \left(\frac{1 - T_{\text{CMB}}(z)}{T_{\text{gas}}(\mathbf{x}, z)} \right)$$



from x_{HII} From ρ_α From ρ_h

Simulating the
Epoch of Reionization

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Halo growth
Implementation

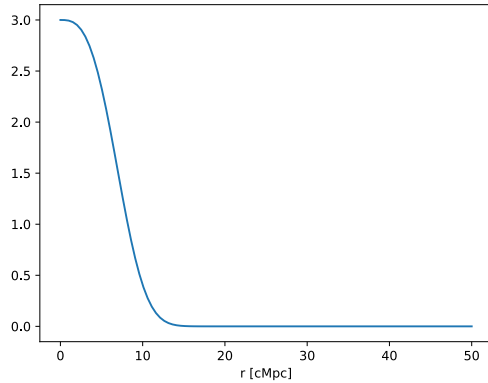
Results

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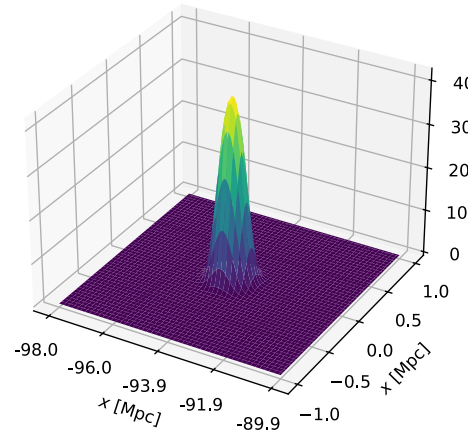
Bibliography

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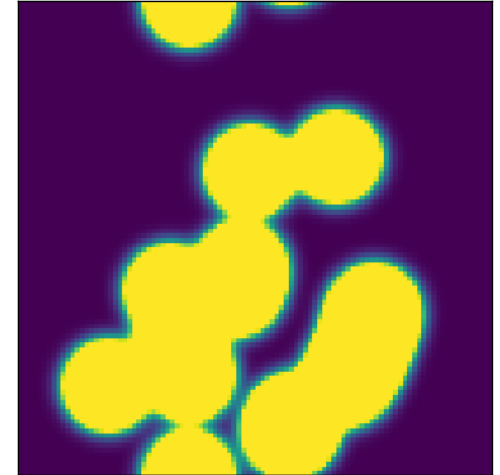
- Procedure
- Postprocessing



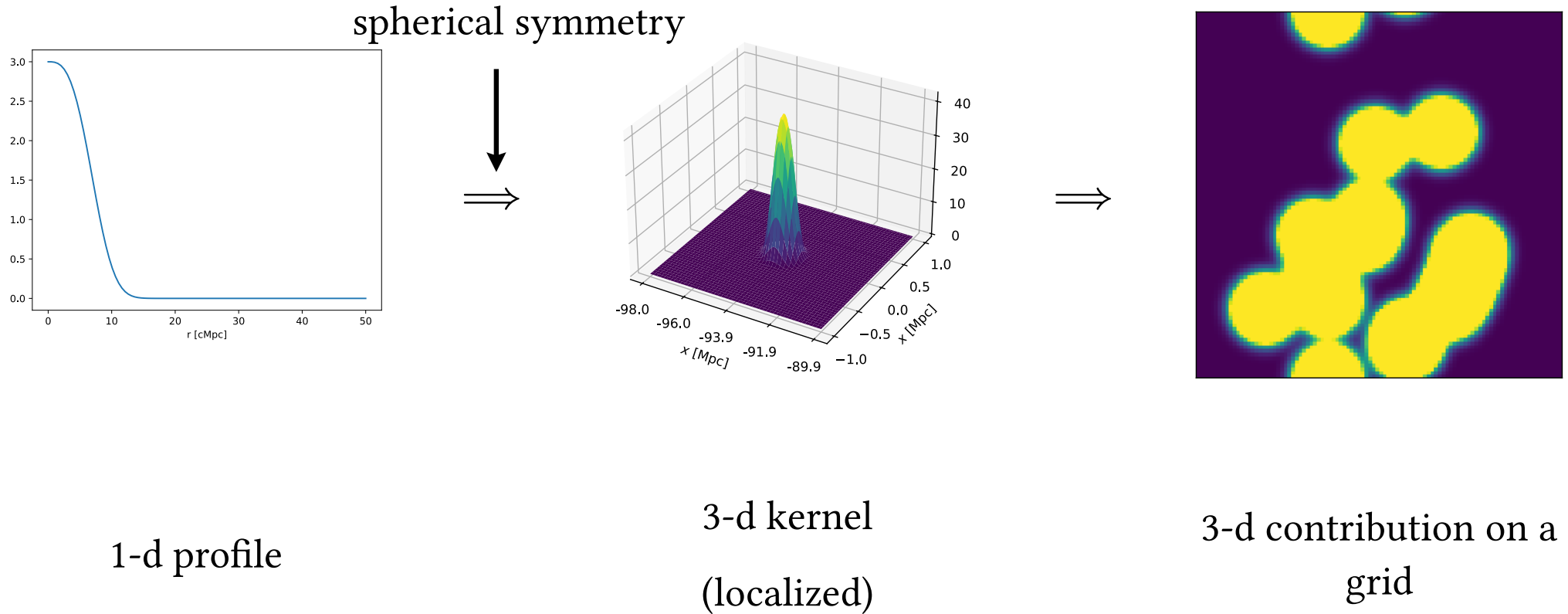
1-d profile

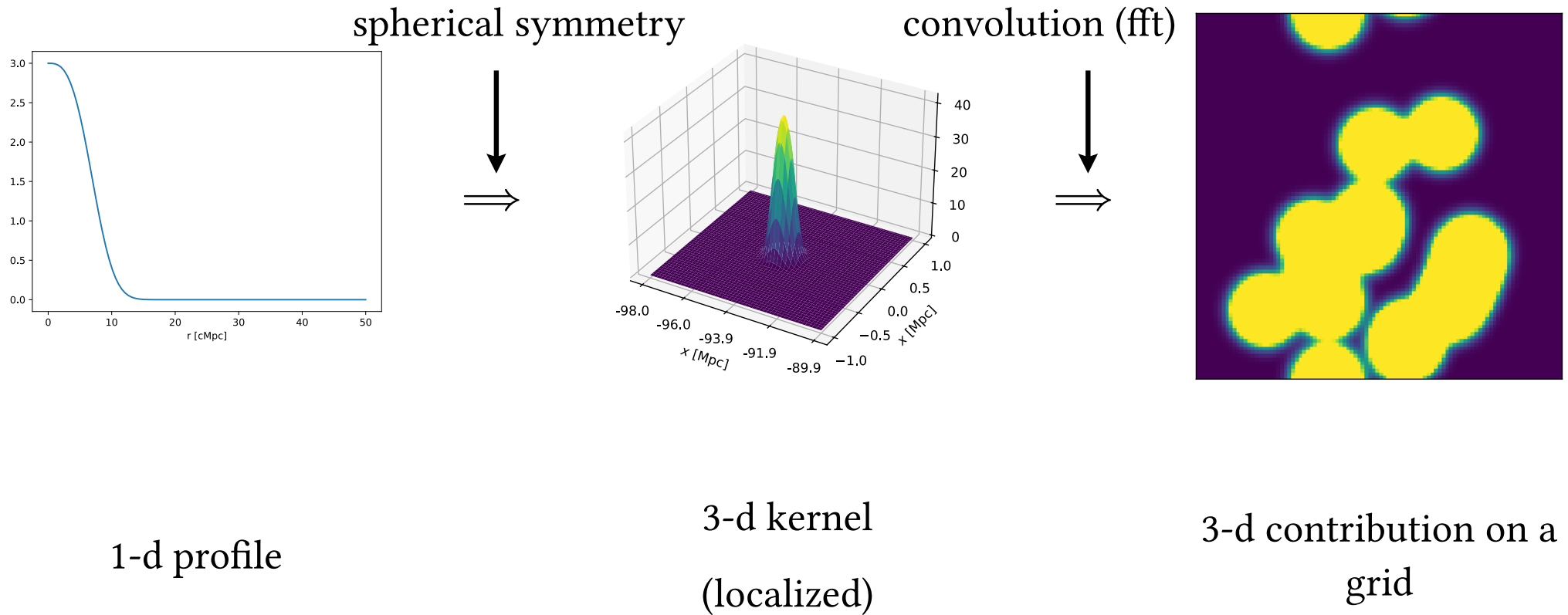


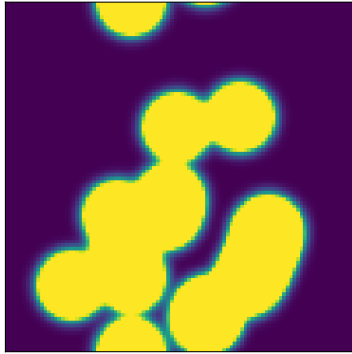
3-d kernel
(localized)



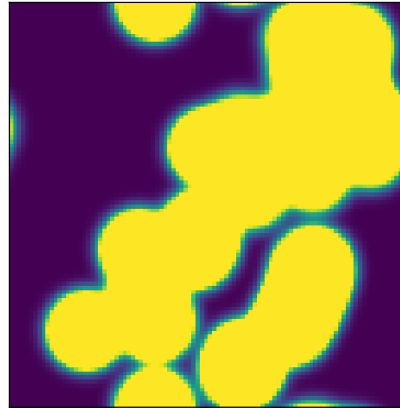
3-d contribution on a
grid







Multiple contributions $\Rightarrow$



$\Rightarrow$ *Postprocessing*
(overlaps, normalization, ...)

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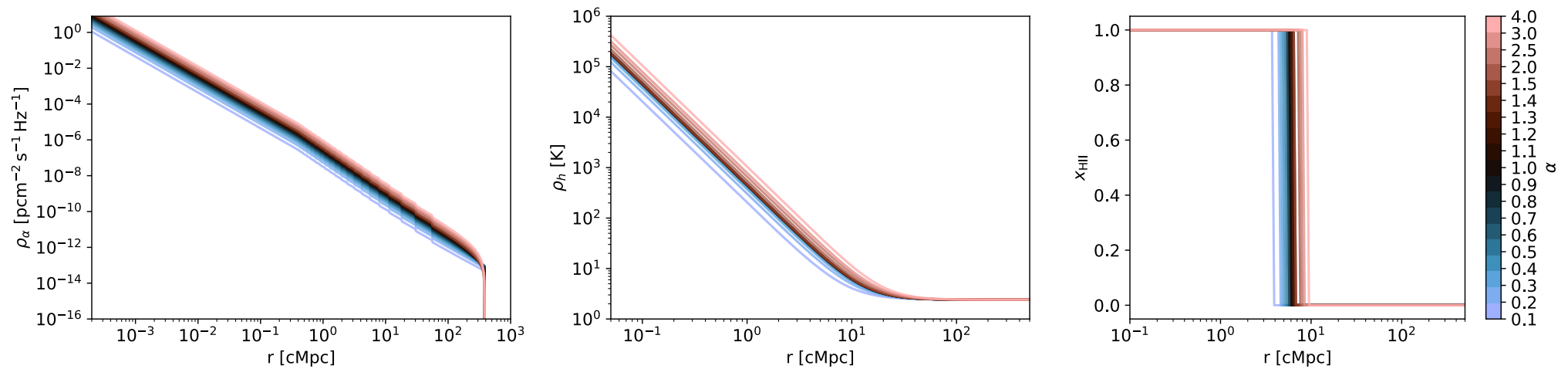
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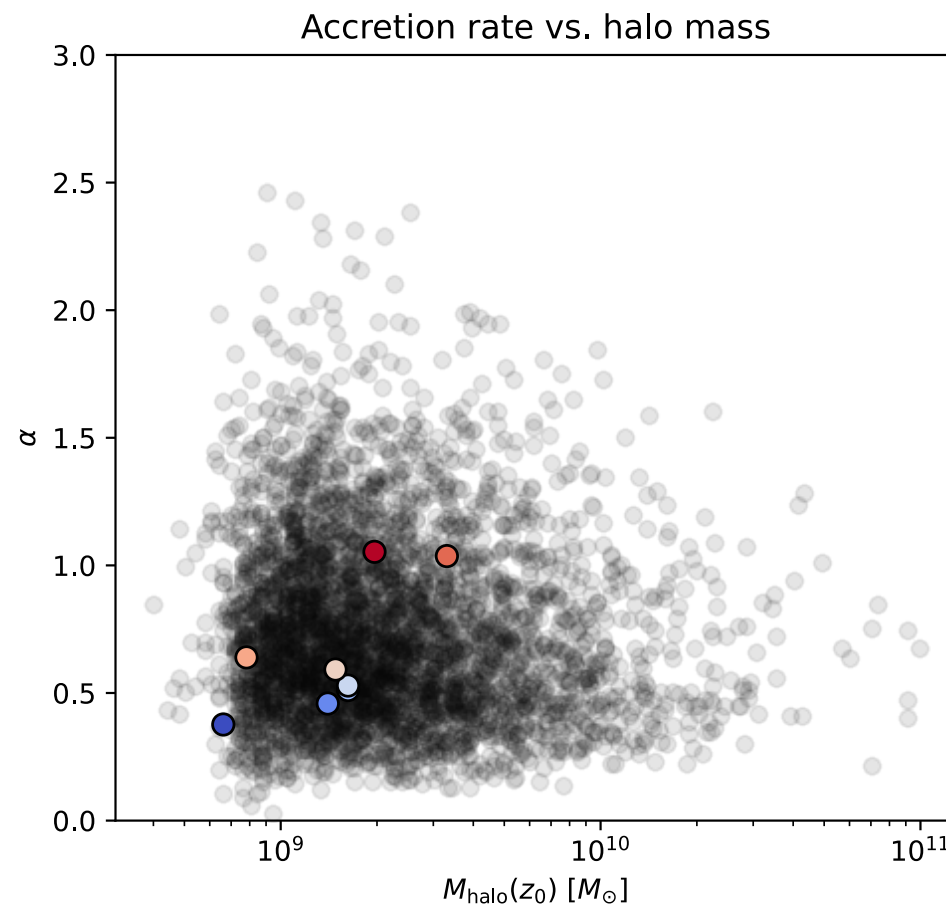
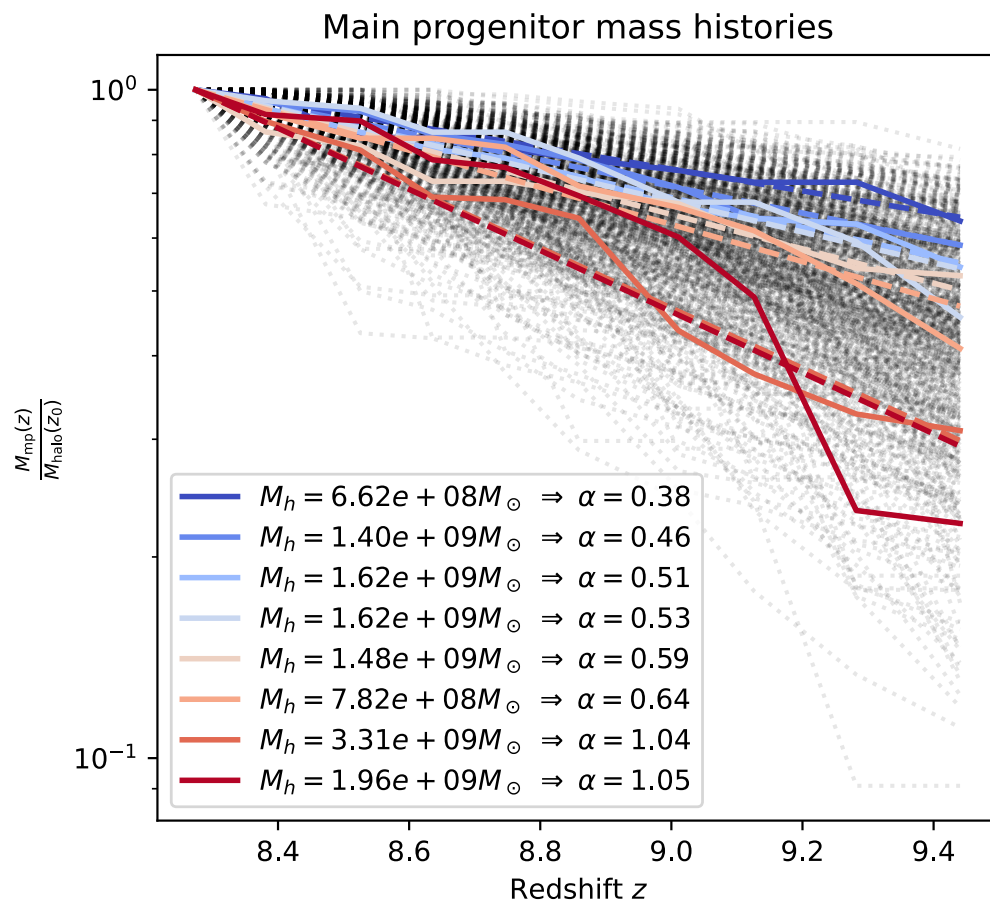
Bibliography

Halo growth

- Effect on the flux profiles
- Inferring growth from THESAN data



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validation

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Implementation

— Simplified usage

 /beorn.py

```
from pathlib import Path
import beorn

current_directory = Path(".")
## Setup the parameters
parameters_file = current_directory / "parameters.yaml"
parameters = beorn.structs.Parameters.from_yaml(parameters_file)
# sample format:
# parameters.solver.redshifts = [6, 20]
# parameters.simulation.file_root = ... / "Thesan-Dark-1"

## Handling of the io
```

```
# this will interface with the input simulation
loader = beorn.load_input_data.ThesanLoader(
    parameters,
    is_high_res = True
)

cache_handler = beorn.io.Handler(current_directory / "cache")
output_handler = beorn.io.Handler(current_directory / "output")
# handlers can also manage logs for us:
# output_handler.save_logs(parameters)

## Computation of the radiation profiles
solver = beorn.radiation_profiles.ProfileSolver(parameters)
profiles = solver.solve()
```

```
## Full 3D painting of the radiation profiles over the specified redshifts
painter = beorn.painting.Painter(
    parameters,
    loader = loader,
    cache_handler = cache_handler,
    output_handler = output_handler
)

grid = painter.paint_full(profiles)
```

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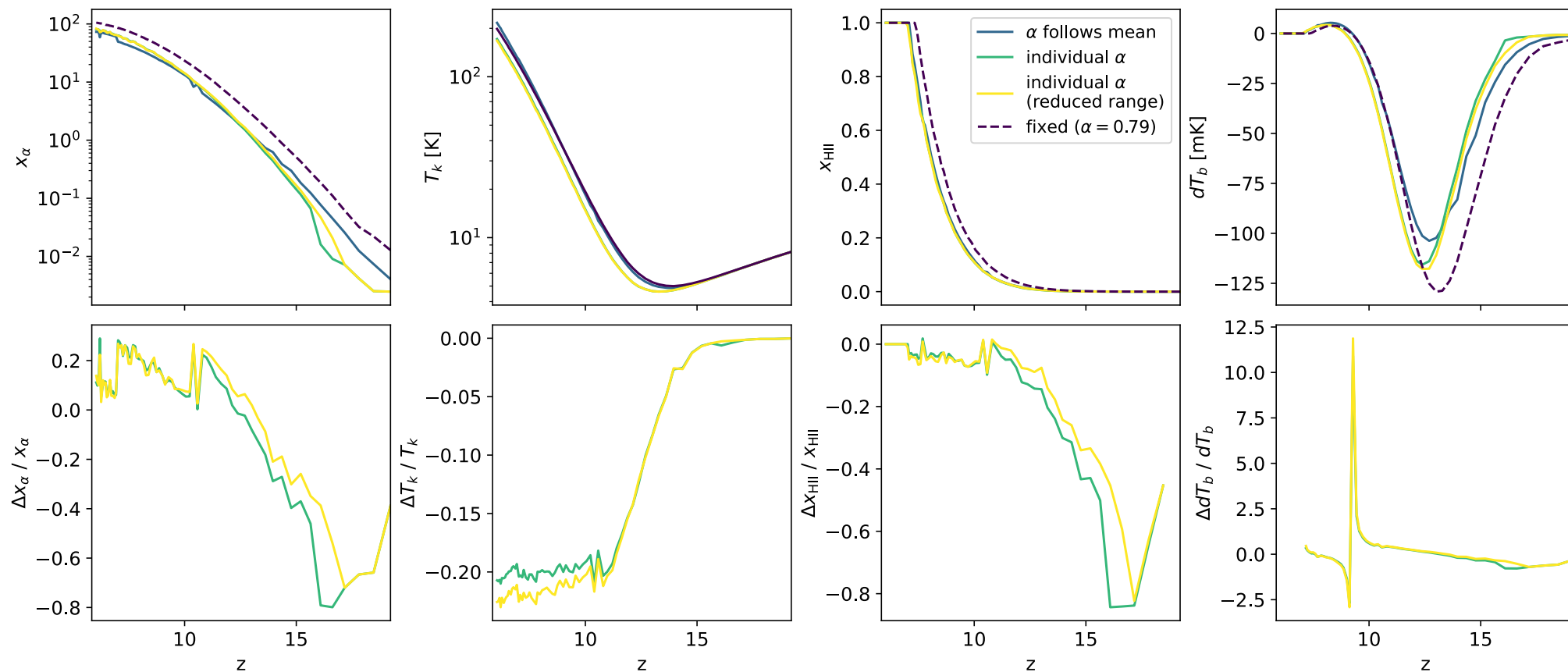
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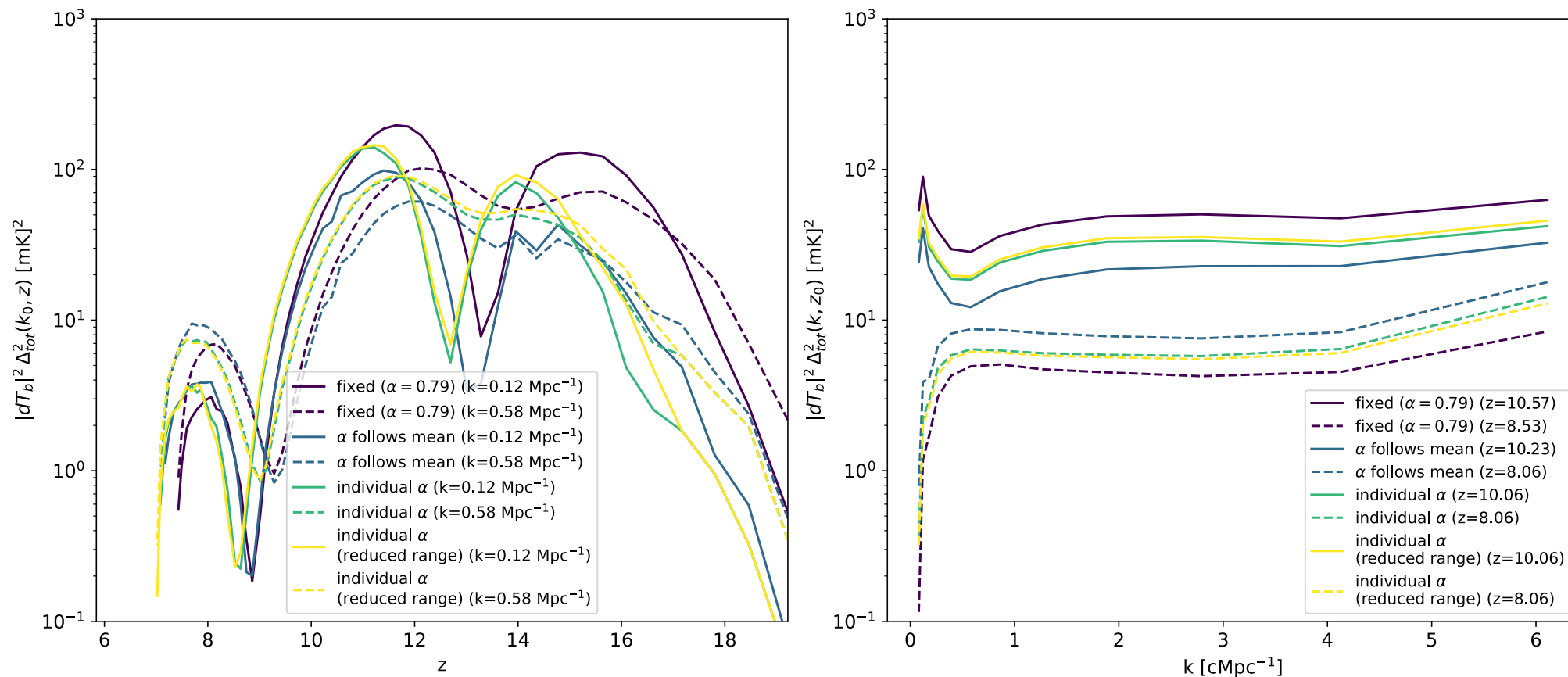
End

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- Map outputs
- Signals





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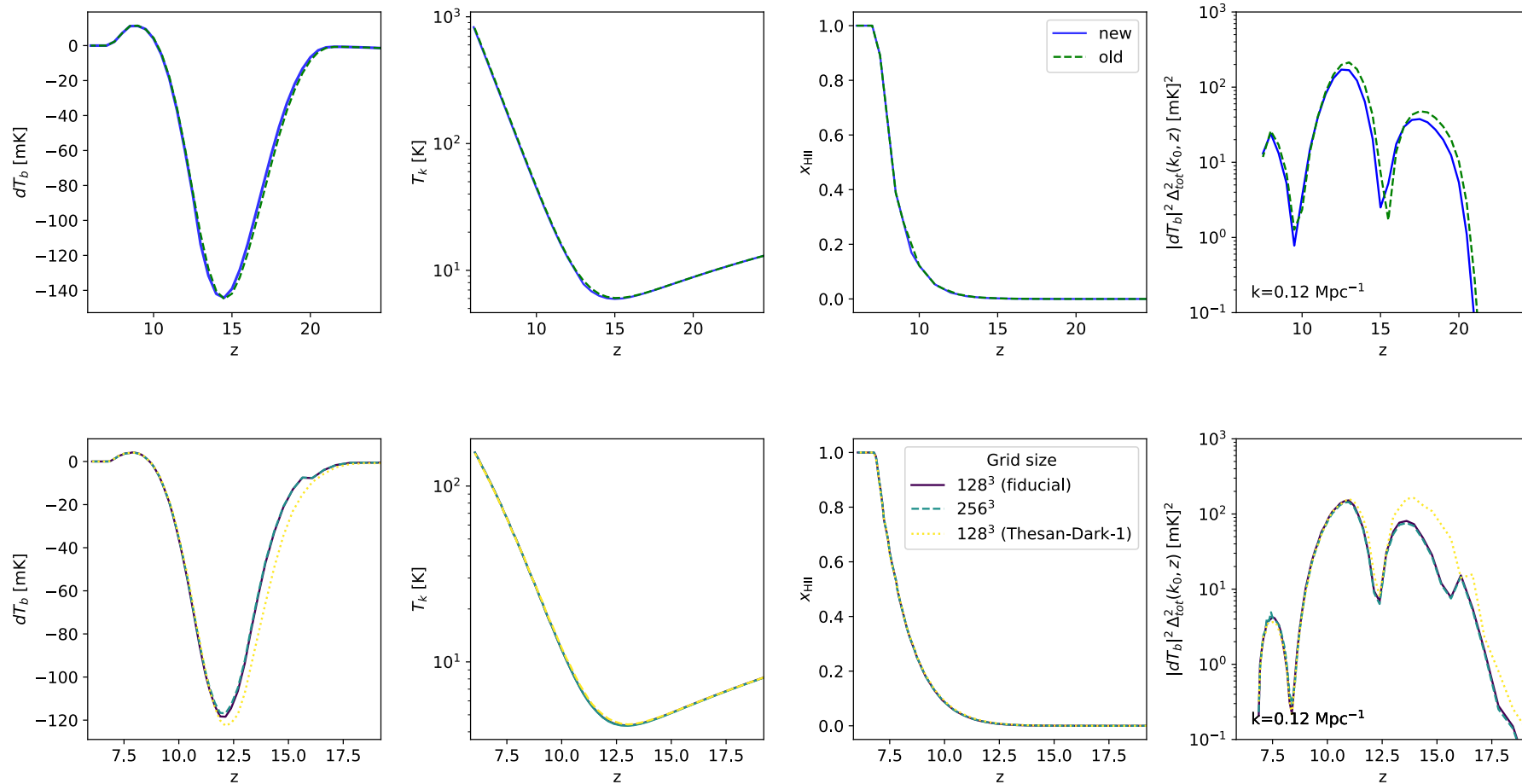
End

— Thank you for your attention

Bibliography

- [1] “CMB spectrum.” [Online]. Available: <https://commons.wikimedia.org/wiki/File:Cmbr.svg>
- [2] J. R. Pritchard and A. Loeb, “21 cm cosmology in the 21st century,” *Reports on Progress in Physics*, vol. 75, no. 8, p. 86901, Aug. 2012, doi: 10.1088/0034-4885/75/8/086901.
- [3] T. Schaeffer, S. K. Giri, and A. Schneider, “<scp>beorn</scp>: a fast and flexible framework to simulate the epoch of reionization and cosmic dawn,” *Monthly Notices of the Royal Astronomical Society*, vol. 526, no. 2, pp. 2942–2959, Sep. 2023, doi: 10.1093/mnras/stad2937.
- [4] A. Schneider, S. K. Giri, and J. Mirocha, “Halo model approach for the 21-cm power spectrum at cosmic dawn,” *Physical Review D*, vol. 103, no. 8, Apr. 2021, doi: 10.1103/physrevd.103.083025.

- A. Schneider, T. Schaeffer, and S. K. Giri, “Cosmological forecast of the 21-cm power
[5] spectrum using the halo model of reionization.” [Online]. Available: <https://arxiv.org/abs/2302.06626>



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