

Simulating the EOR with self-consistent growth of galaxies

Master's thesis presentation

ETH Zürich, University of Zürich

Simulating the Epoch of Reionization

Bibliography

Hello

Simulating the Epoch of Reionization

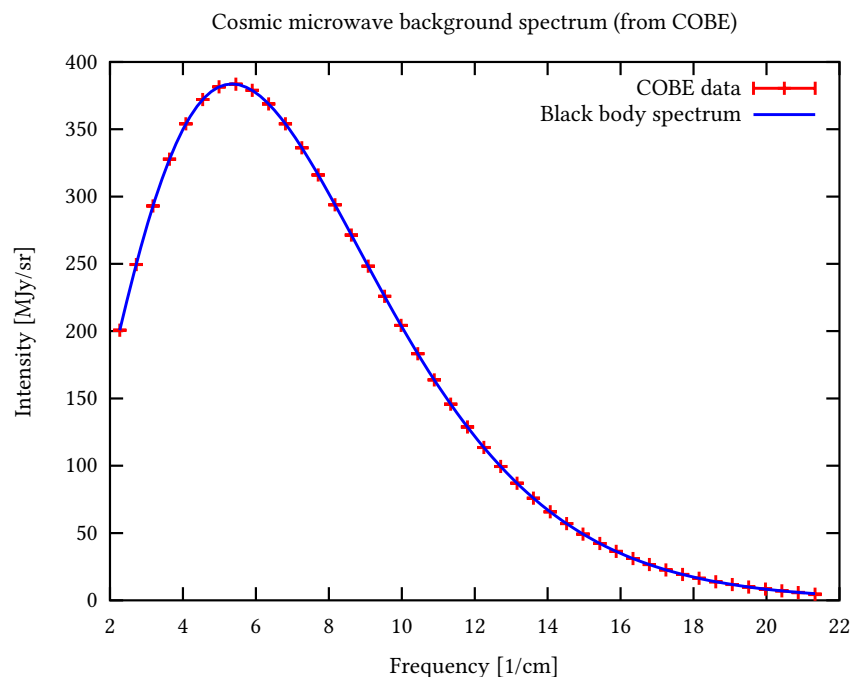
- The Epoch of Reionization
- The 21cm signal
- The halo model of reionization
- The current state of simulations
- BEO RN

- Marks the universe's last major phase transition: from neutral to ionized hydrogen.
- Shapes the large-scale structure of the intergalactic medium (IGM).
- Is strongly linked to the formation and growth of the first galaxies and black holes.
- Sets the stage for many observables: CMB secondary anisotropies, 21-cm signal, high- z galaxy surveys.

The 21cm signal

3 / 16

The brightness temperature describes the difference between the CMB temperature and the spin temperature of neutral hydrogen

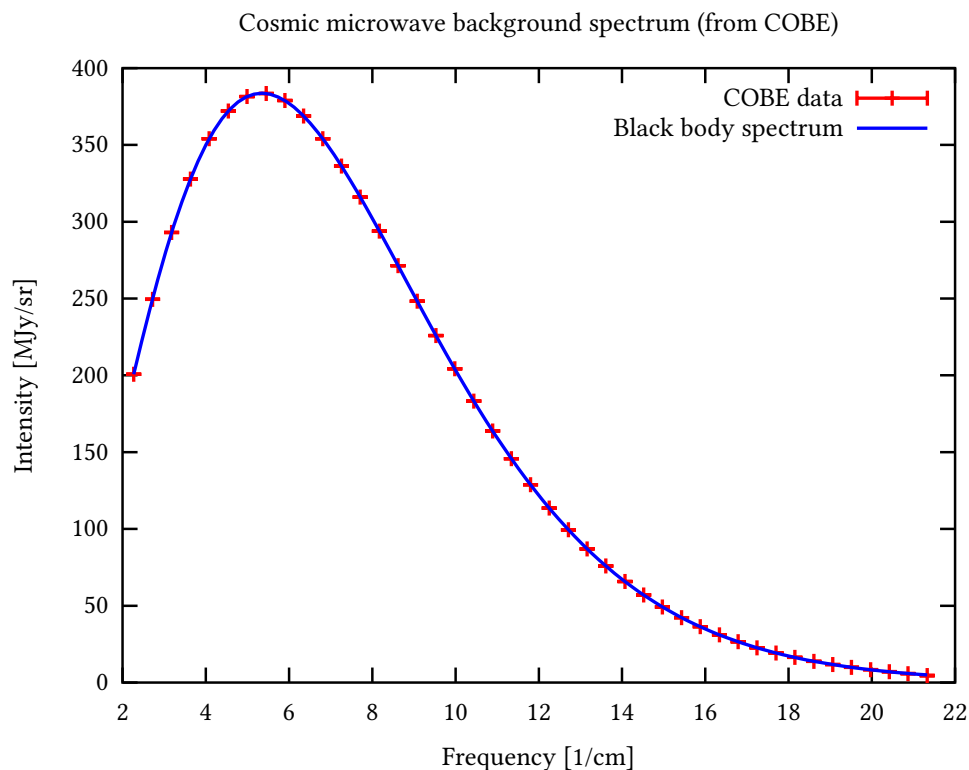


[1]

The 21cm signal

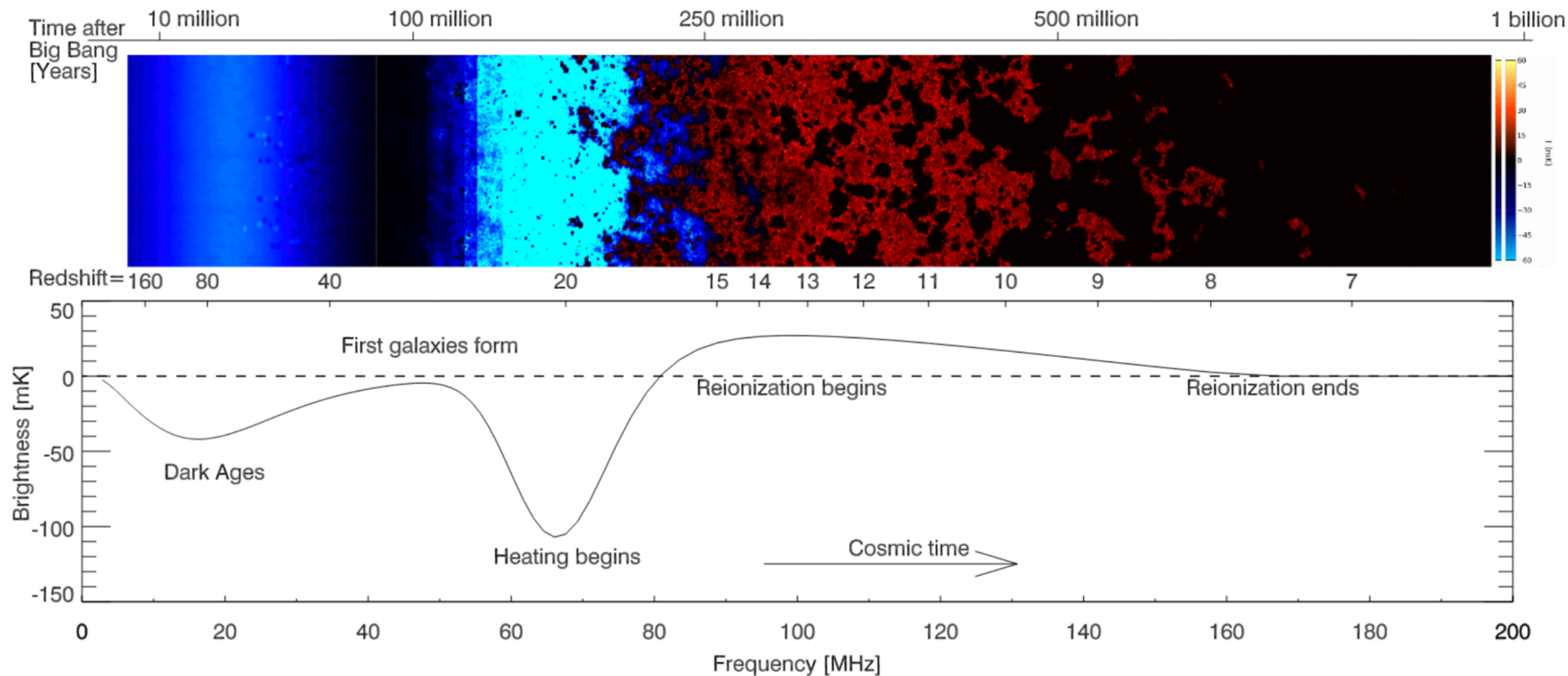
4 / 16

Removing the contribution from the black body spectrum of the CMB yields the explicit 21-cm signal:



The 21cm signal

5 / 16



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[2]

$$dT_b(\mathbf{x}, z) \simeq T_0(z) \cdot x_{\text{HI}}(\mathbf{x}, z) \cdot (1 + \delta_b(\mathbf{x}, z)) \cdot \frac{x_\alpha(\mathbf{x}, z)}{1 + x_\alpha(\mathbf{x}, z)} \cdot \left(\frac{1 - T_{\text{CMB}}(z)}{T_{\text{gas}}(\mathbf{x}, z)} \right)$$

(e.g [2])

Following [3], [4]:

$$\rho_{\alpha}(r \mid M, z) = \frac{(1+z)^2}{4\pi r^2} \cdot \sum_{n=2}^{n_m} f_n \cdot \varepsilon_{\alpha}(\nu') \cdot f_{\star} \cdot \dot{M}(z' \mid M, z)$$

$$\frac{3}{2} \cdot \frac{d\rho_h(r \mid M, z)}{dz} = \frac{3\rho_h(r \mid M, z)}{1+z} - \frac{\rho_{\text{xray}}(r \mid M, z)}{k_B(1+z)H(z)}$$

$$x_{\text{HII}}(r \mid M, z) = \theta_{\text{H}} \left[R_b(M, z) - r \right]$$

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- slow and big
- radiative transfer

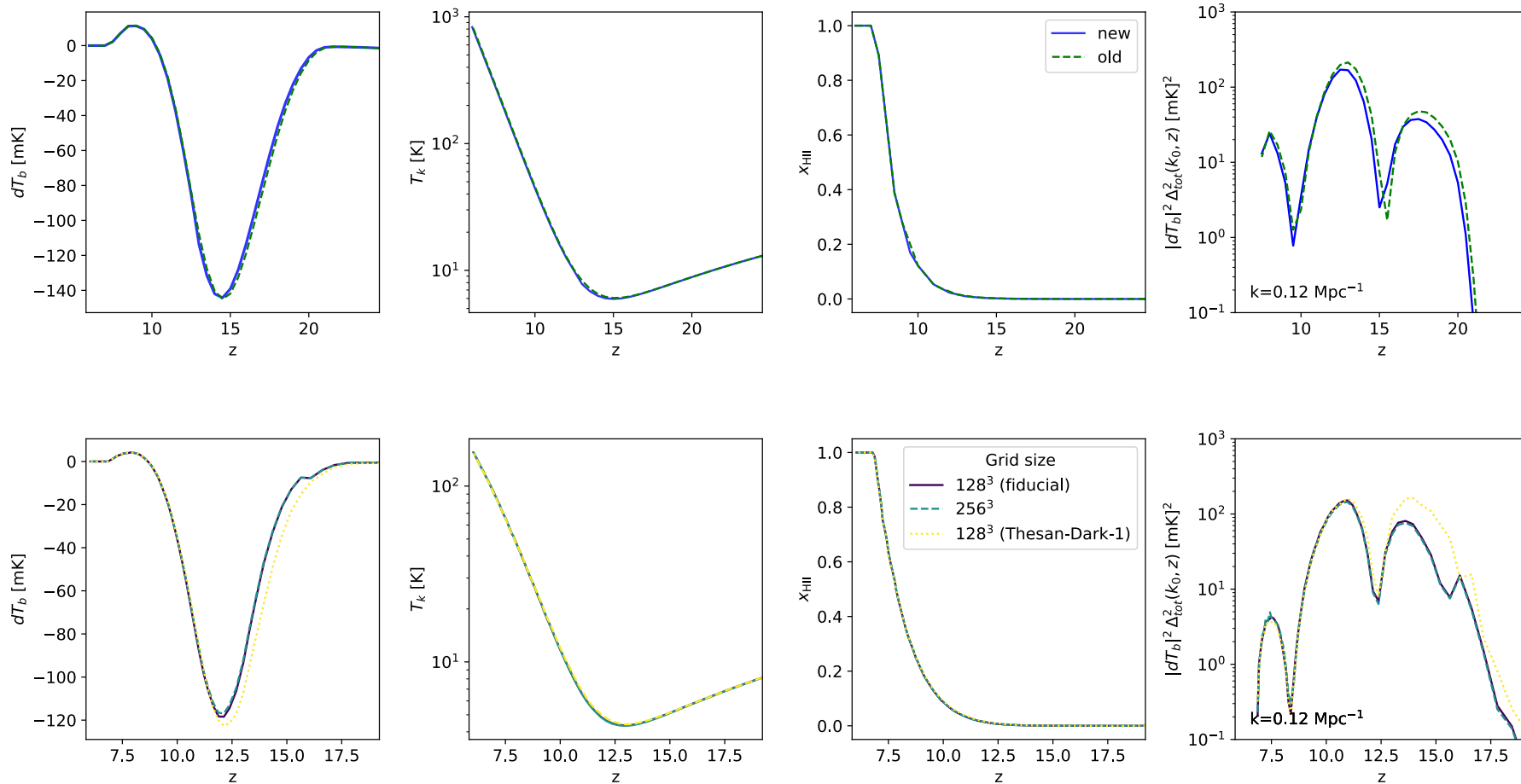
⇒ semi-numerical approaches such as BEO RN [5]

validation

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Hello

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— This works right

This works right

16 / 16

$$E = mc^2$$

This works right

16 / 16

$$E = mc^2$$

$$F = ma$$